

Asymptotic Safety for Gravity

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1 ERGE

2 Fixed Point

ERGE

We want to solve the ERGE equation

$$\partial_t \Gamma_k = \frac{1}{2} \text{Tr} \left[(\Gamma_k^2 + \mathcal{R}_k)^{-1} \partial_t \mathcal{R}_k \right] \quad (1)$$

for gravity. Where $t = \log(k/k_0)$ for some fixed k_0 .

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Γ_k admits an expansion of the form

$$\Gamma_k[g_{\mu\nu}, g_i^{(2n)}] = \sum_{n=0}^{\infty} \sum_i g_i^{(2n)} \mathcal{P}_i^{(2n)}(g_{\mu\nu}) \quad (2)$$

where $g_i^{(2n)}$ are the coupling constants and $\mathcal{P}_i^{(2n)}(g_{\mu\nu})$ are all possible operators constructed from $g_{\mu\nu}$ and its n -derivates, which are compatible with the symmetry of the theory.

With the equations (1) and (2) we have

$$\frac{1}{2} \text{Tr} \left[(\Gamma_k^2 + \mathcal{R}_k)^{-1} \partial_t \mathcal{R}_k \right] = \sum_{n=0}^{\infty} \sum_i \beta_i^{(2n)} \mathcal{P}_i^{(2n)}(g_{\mu\nu}) \quad (3)$$

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To solve this equation we need an specific form of the cutoff \mathcal{R}_k and Γ_k . For the scalar theory with \mathbb{Z}_2 symmetry we have

$$\Gamma_k^{(2)}(x, y) = (-\partial_x^2 + 2V'_k + 4\phi V''_k) \delta(x - y) \quad (4)$$

with $V_k = V_k(\phi^2)$ and primes denote ϕ^2 derivatives.

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Observe that we have an expresion of the form

$$\Gamma_k^{(2)} = -\partial^2 + \mathbf{E} \quad (5)$$

Gravity

In general the inverse of $\Gamma^{(2)}$ is an operator of the form

$$\Delta = -\nabla^2 + \mathbf{E} \quad (6)$$

where $\mathbf{E} = \mathbf{E}_1 + \mathbf{E}_2$ is an operator acting on the fields. \mathbf{E}_1 doesn't contain couplings. \mathbf{E}_2 contain couplings.

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We consider that $\mathbf{E}_2 = 0$. This is called a cutoff of type II. Therefore the inverse propagator is

$$P_k(\Delta) = \Delta + R_k(\Delta) \quad (7)$$

The expansion of the effective action for pure gravity is

$$\Gamma_k^{(n \leq 2)} = \int d^4x \sqrt{g} \sum_{n=0}^2 \sum_i g_i^{(2n)} \mathcal{M}_i^{(2n)}$$

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$$\Gamma_k^{(n \leq 2)} = \int d^4x \sqrt{g} \sum_{n=0}^2 \sum_i g_i^{(2n)} \mathcal{M}_i^{(2n)}$$

with $g^{(0)} = 2Z\Lambda$, $g^{(2)} = -Z$, $g_1^{(4)} = 1/(2\lambda)$, and $g_2^{(4)} = 1/\xi$.

$$\begin{aligned} \mathcal{M}^{(0)} &= 1 & \mathcal{M}^{(2)} &= R \\ \mathcal{M}_1^{(4)} &= C^2 & \mathcal{M}_2^{(4)} &= R^2 \end{aligned}$$

R is the scalar curvature and C^2 is the square of the Weyl tensor.

Therefore we get

$$\begin{aligned}\Gamma_k^{(n \leq 2)} &= \int d^4x \sqrt{g} \left[2Z\Lambda - ZR + \frac{1}{2\lambda} C^2 + \frac{1}{\xi} R^2 \right] \\ &:= (\Gamma_k)_{grav}\end{aligned}\tag{8}$$

for pure gravity.

Now, consider n_s scalar fields, n_D Dirac fields and n_M Maxwell fields minimally coupled to gravity, the effective action is

$$\begin{aligned}\Gamma_k &= \int d^4x \sqrt{g} \left[\frac{1}{2} \nabla_\mu \phi \nabla^\mu \phi + \bar{\psi} \gamma^\mu \nabla_\mu \psi + \frac{1}{4} F^{\mu\nu} F_{\mu\nu} \right] \\ &:= (\Gamma_k)_{\text{matt}}\end{aligned}\tag{9}$$

where ∇ is the covariant derivative.

Hence, the effective action is given by

$$\Gamma_k(g_{\mu\nu}, \phi, \psi, A_\mu) = (\Gamma_k)_{grav} + (\Gamma_k)_{Matt} \quad (10)$$

Pag. [15]

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$$(\Gamma_k^{(2)} + R_k)^{-1} = ((\Gamma_k^{(2)})_{\text{Matt}} + R_k)^{-1}$$

where

$$\begin{aligned} (\Gamma_k^{(2)})_{\text{Matt}}(x, y) &= n_s \left(\frac{\delta^2(\Gamma_k)_{\text{Matt}}}{\delta\phi(x)\delta\phi(y)} \right) + n_D \left(\frac{\delta^2(\Gamma_k)_{\text{Matt}}}{\delta\bar{\psi}(x)\delta\psi(y)} \right) \\ &\quad + n_M \left(\frac{\delta^2(\Gamma_k)_{\text{Matt}}}{\delta A_\mu(x)\delta A_\nu(y)} \right) \end{aligned}$$

We obtain the ERGE:

ERGE

$$\begin{aligned} \partial_t \Gamma_k = & \frac{n_S}{2} \text{Tr}_{(S)} \left(\frac{\partial_t R_k(\Delta^{(S)})}{P_k(\Delta^{(S)})} \right) - \frac{n_D}{2} \text{Tr}_{(D)} \left(\frac{\partial_t R_k(\Delta^{(D)})}{P_k(\Delta^{(D)})} \right) \\ & + \frac{n_M}{2} \text{Tr}_{(M)} \left(\frac{\partial_t R_k(\Delta^{(M)})}{P_k(\Delta^{(M)})} \right) - \frac{n_{gh}}{2} \text{Tr}_{(gh)} \left(\frac{\partial_t R_k(\Delta^{(gh)})}{P_k(\Delta^{(gh)})} \right) \end{aligned} \quad (11)$$

with a type II cutoff $P_k(\Delta^{(A)}) = \Delta^{(A)} + R_k(\Delta^{(A)})$ we obtain for each case

scalar

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Ghost

$$\Delta^{(gh)} = -\nabla^2 \quad (15)$$

Hence

Trace

$$\begin{aligned}\partial_t \Gamma_k &= \frac{1}{2} \frac{1}{(4\pi)^2} \int d^4x \sqrt{g} \left[(n_s - 4n_D + 2n_M) Q_2 \left(\frac{\partial_t R_k}{P_k} \right) \right. \\ &\quad + \frac{1}{6} R (n_s - 2n_D - 4n_M) Q_1 \left(\frac{\partial_t R_k}{P_k} \right) \\ &\quad + \frac{1}{180} \left((3n_s + 18n_D + 36n_M) C^2 - (n_s + 11n_D + 62n_M) E \right. \\ &\quad \left. \left. + 5n_s R^2 + 12(n_s + n_D - 3n_M) \nabla^2 R \right) + \dots \right] \quad (16)\end{aligned}$$

C^2 is the square of the Weyl's tensor and

$$E = R_{\mu\nu\rho\sigma} R^{\mu\nu\rho\sigma} - 4R_{\mu\nu} R^{\mu\nu} + R^2. \text{Pag. [15]}$$

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Remark

The coefficients of the 4-derivative terms are scheme independent.

The function Q is defined as

$$Q_n(W) = \frac{1}{\Gamma(n)} \int_0^\infty dz z^{n-1} W(z)$$

Fixed Point

In order to look for a fixed point we have to obtain the equations for the beta functions. This is done with (10) and (16). First we note that

$$\begin{aligned}\partial_t(\Gamma_k)_{grav} &= \int d^4x \sqrt{g} \left[\partial_t(2Z\Lambda) + \partial_t(-Z)R + \partial_t\left(\frac{1}{2\lambda}\right)C^2 \right. \\ &\quad \left. + \partial_t\left(\frac{1}{\xi}\right)R^2 \right] \\ &= \int d^4x \sqrt{g} \left[\partial_t(g^{(0)}) + \partial_t(g^{(2)})R + \partial_t(g_1^{(4)})C^2 \right. \\ &\quad \left. + \partial_t(g_2^{(4)})R^2 \right]\end{aligned}$$

Then

Equations

$$\partial_t g^{(0)} = \frac{1}{2(4\pi)^2} (n_s - 4n_D + 2n_M) Q_2 \left(\frac{\partial_t R_k}{P_k} \right) \quad (17)$$

$$\partial_t g^{(2)} = \frac{1}{12(4\pi)^2} (n_s - 2n_D - 4n_M) Q_1 \left(\frac{\partial_t R_k}{P_k} \right) \quad (18)$$

with the optimized cutoff

$$R_k(z) = (k^2 - z)\theta(k^2 - z)$$

The integrals Q are

$$Q_n \left(\frac{\partial_t R_k}{(P_k + q)^l} \right) = \frac{2}{n!} \frac{1}{(1 + \tilde{q})^l} k^{2(n-l+1)} \quad (19)$$

with $\tilde{q} = k^{-2}q$.

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with $\tilde{q} = k^{-2}q$. Hence

$$Q_2 \left(\frac{\partial_t R_k}{P_k} \right) = k^4 \quad (20)$$

$$Q_1 \left(\frac{\partial_t R_k}{P_k} \right) = 2k^2 \quad (21)$$

Then

$$\partial_t g^{(0)} = a^{(0)} k^4 \quad (22)$$

$$\partial_t g^{(2)} = a^{(0)} k^2 \quad (23)$$

with

$$a^{(0)} = \frac{n_s - 4n_D + 2n_M}{2(4\pi)^2}$$

$$a^{(2)} = \frac{n_s - 2n_D - 4n_M}{6(4\pi)^2}$$

In general $\partial_t g^{(n)} = a^{(n)} k^{4-n}$.

With $\tilde{g}^{(n)} = k^{n-4} g^{(n)}$

$$\partial_t \tilde{g}^{(n)} = (n-4) \tilde{g}^{(n)} + k^{n-4} \partial_t g^{(n)} = (n-4) \tilde{g}^{(n)} + a^{(n)}$$

and

$$Z = \frac{1}{16\pi G}$$

we have

$$\begin{aligned} \partial_t \tilde{g}^{(0)} &= \partial_t (2k^{-4} Z \Lambda) = \partial_t \left(\frac{\tilde{\Lambda}}{8\pi \tilde{G}} \right) \\ &= \frac{\partial_t \tilde{\Lambda}}{8\pi \tilde{G}} - \frac{\tilde{\Lambda}}{8\pi \tilde{G}^2} \partial_t \tilde{G} \\ &= -\frac{\tilde{\Lambda}}{2\pi \tilde{G}} + a^{(0)} \end{aligned}$$

With this we get

$$(\partial_t \tilde{\Lambda}) \tilde{G} - (\partial_t \tilde{G}) \tilde{\Lambda} = -4 \tilde{G} \tilde{\Lambda} + 8\pi \tilde{G}^2 a^{(0)}$$

in the same way we obtain with $g^{(2)}$

$$\partial_t \tilde{G} = 2 \tilde{G} + 16\pi \tilde{G}^2 a^{(2)}$$

with these equations finally

$$\partial_t \tilde{\Lambda} = -2 \tilde{\Lambda} + 16\pi \tilde{G} \tilde{\Lambda} + 8\pi \tilde{G} a^{(0)} \quad (24)$$

$$\partial_t \tilde{G} = 2 \tilde{G} + 16\pi \tilde{G}^2 a^{(2)} \quad (25)$$

From (24) if

$$\partial_t \tilde{G} = 0 = 2\tilde{G}(1 + 8\pi\tilde{G}a^{(2)})$$

hence

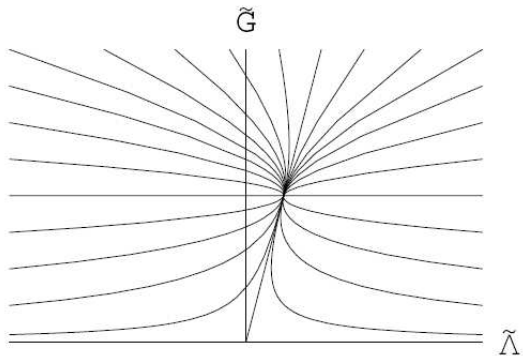
$$\begin{aligned}\tilde{G}_* &= \frac{1}{8\pi a^{(2)}} = \frac{1}{8\pi} \frac{6(4\pi)^2}{(n_s - 2n_D - 4n_M)} \\ \tilde{G}_* &= \frac{12\pi}{n_s - 2n_D - 4n_M}\end{aligned}\tag{26}$$

and

$$\tilde{\lambda}_* = -\frac{3}{4} \frac{n_s - 4n_D + 2n_M}{n_s - 2n_D - 4n_M}\tag{27}$$

There is a non trivial Fixed Point.

The flow is given by



For pure gravity, we consider the Einstein-Hilbert truncation

$$\Gamma_k^{(n \leq 2)} = \int d^4x \sqrt{g} [2Z\Lambda - ZR] + S_{GF} + S_{gh} \quad (28)$$

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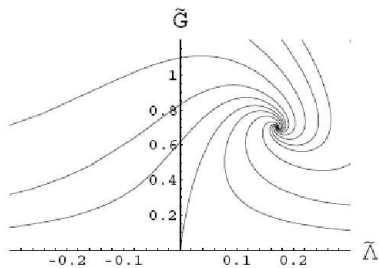
$$\Gamma_k^{(n \leq 2)} = \int d^4x \sqrt{g} [2Z\Lambda - ZR] + S_{GF} + S_{gh} \quad (28)$$

the equations for the beta functions are






$$\partial_t \tilde{\Lambda} = \frac{-2(1 - 2\tilde{\Lambda})^2 \tilde{\Lambda} + \frac{36 - 41\tilde{\Lambda} + 42\tilde{\Lambda}^2 - 600\tilde{\Lambda}^3}{72\pi} \tilde{G} + \frac{467 - 572\tilde{\Lambda}}{288\pi^2} \tilde{G}^2}{(1 - 2\tilde{\Lambda})^2 - \frac{29 - 9\tilde{\Lambda}}{72\pi} \tilde{G}} \quad (29)$$

$$\partial_t \tilde{G} = \frac{2(1 - 2\tilde{\Lambda})^2 \tilde{G} - \frac{373 - 654\tilde{\Lambda} + 600\tilde{\Lambda}^2}{72\pi} \tilde{G}^2}{(1 - 2\tilde{\Lambda})^2 - \frac{29 - 9\tilde{\Lambda}}{72\pi} \tilde{G}} \quad (30)$$

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References

-  R. Percacci (2007). Asymptotic Safety. arXiv:hep-th/0709.3851
-  R. Percacci. Web page.
<http://www.percacci.it/roberto/physics/as/index.html>
-  A. Codello, R. Percacci and C. Rahmede (2008). Investigating the Ultraviolet Properties of Gravity with a Wilsonian Renormalization Group Equation. arXiv:hep-th/2909
-  C. Wetterich (1993). Exact Evolution Equation for the Effective Potential. Phys. Lett. B301, 90.
-  S. Weinberg (1979). Ultraviolet Divergences in Quantum Theories of Gravitation. In general Relativity: An Einstein centenary survey, ed. S. Hawking and W. Israel, chapter 16, pp. 790-831; Cambridge University Press.