General Boundary Quantum Field Theory in curved space

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General Boundary Formulation

The GBF is an **axiomatic** formulation of quantum theory which combines

 the mathematical framework of Topological Quantum Field Theory (association of algebraic structures to geometric ones) with

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a generalization of the Born's rule to extract probabilities.

Basic structures

In the GBF algebraic structures are associated to geometric ones.

Geometric structures (representing pieces of spacetime):

- **b** hypersurfaces: oriented manifolds of dimension d-1
- **regions**: oriented manifolds of dimension d with boundary

Algebraic structures:

- To each hypersurface Σ associate a Hilbert space \mathcal{H}_{Σ} of states.
- ► To each region M with boundary ∂M associate a linear amplitude map $\rho_M : \mathcal{H}_{\partial M} \to \mathbb{C}$
- As in AQFT, observables are associated to spacetime regions: An observable O in a region M is a linear map ρ^O_M : H_{∂M} → C, called observable map.

We want to construct a general boundary quantum field theory in curved space.



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Classical Theory

We consider the linear theory of a massive Klein-Gordon field in a 4d Lorentzian spacetime $ds^2 = g_{\mu\nu}dx^{\mu}dx^{\nu}$.

The action of field in a spacetime region M is

$$S_{M,0}[\phi] = \frac{1}{2} \int_{M} \mathrm{d}^{4} x \sqrt{-g} \left(g^{\mu\nu} \partial_{\mu} \phi \partial_{\nu} \phi - m^{2} \phi^{2} \right)$$

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- \blacktriangleright the integration is extended over the sapcetime region M
- g is determinant of the metric $g_{\mu\nu}$
- we used the notation $\partial_{\mu} = \partial/\partial x^{\mu}$

Assumptions

We assume the following:

- ► The spacetime region M admits a **foliation** in terms of hypersurfaces, *not necessarily spacelike*, described in terms of a smooth coordinate system (τ, \underline{x}) . The coordinates on the leaves of the foliation are denoted by $\underline{x} = (x^1, x^2, x^3) \in I \subset \mathbb{R}^3$ and the leaves are parametrized by the coordinate $\tau \in J \subset \mathbb{R}$. Notice that τ and \underline{x} are not required to be timelike and spacelike coordinates respectively.
- The metric takes a block-diagonal form with respect to the coordinates (τ, <u>x</u>), i.e. g^{τxⁱ} = 0 = g^{xⁱτ} for all i ∈ {1,2,3}.
- Let L_Σ be the space of solutions of the e.o.m. in a neighborhood of a hypersurface Σ of the foliation. We define the symplectic potential as the one-form on L_Σ,

$$[\xi,\xi'] := -\int_{\Sigma} \mathrm{d}^3\underline{x}\,\sqrt{g^{(3)}}\,\xi(\vartheta_\tau\xi'),$$

where $g^{(3)}$ is the determinant of the induced 3-metric on the hypersurface Σ , and $\xi, \xi' \in L_{\Sigma}$. The corresponding **symplectic form** is

$$\omega_{\Sigma}(\xi,\xi') = \frac{1}{2}[\xi,\xi'] - \frac{1}{2}[\xi',\xi] \qquad \forall \xi,\xi' \in L_{\Sigma}.$$

The Euler-Lagrangian equations are be solved by the method of separation of variables, and a solution takes the form

$$\Phi(\tau,\underline{x}) = \int \mathrm{d}\underline{k} \left(d_{\mathsf{a}}(\underline{k}) X_{\mathsf{a},\underline{k}}(\tau) Y_{\mathsf{a},\underline{k}}(\underline{x}) + d_{\mathsf{b}}(\underline{k}) X_{\mathsf{b},\underline{k}}(\tau) Y_{\mathsf{b},\underline{k}}(\underline{x}) \right),$$

where X_a and X_b are the two independent solutions of the part of the equation of motion depending only on the τ variable.

Solutions can also be written as

$$\phi(\tau,\underline{x}) = (X_{a}(\tau)Y_{a})(\underline{x}) + (X_{b}(\tau)Y_{b})(\underline{x}),$$

where now $X_{a,b}$ are understood as operators from the space of initial data $Y_{a,b}$ to solution on Σ_{τ} . We assume that $X_{a,b}$ commute with each other and are **invertible**.

The most studied quantum field theories on curved spaces are encompassed by these assumptions.

Slice region

- ► The **slice region** *M* is defined as the spacetime region is bounded by the disjoint union of two constant- τ hypersurfaces, $\Sigma_1 = \{(\tau, \underline{x}) : \tau = \tau_1\}$ and $\Sigma_2 = \{(\tau, \underline{x}) : \tau = \tau_2\}$, namely $M = [\tau_1, \tau_2] \times \mathbb{R}^3$.
- Let $\varphi_1(\underline{x}) := \varphi(\tau_1, \underline{x})$ and $\varphi_2(\underline{x}) := \varphi(\tau_2, \underline{x})$ be the boundary field configurations on Σ_1 and Σ_2 .
- The classical solution can be expressed as

$$\phi(\tau,\underline{x}) = \left(\frac{\Delta(\tau,\tau_2)}{\Delta(\tau_1,\tau_2)}\,\varphi_1\right)(\underline{x}) + \left(\frac{\Delta(\tau_1,\tau)}{\Delta(\tau_1,\tau_2)}\,\varphi_2\right)(\underline{x}),$$

where $\Delta(\tau_1, \tau_2) := X_a(\tau_1)X_b(\tau_2) - X_a(\tau_2)X_b(\tau_1)$. All the deltas must be understood as operators acting on the boundary field configurations φ_1 and φ_2 , and we assume that these operators are invertible.

> The free action in the slice region takes the form

$$\begin{split} \mathcal{S}_{[\tau_1,\tau_2],0}(\varphi) &= \frac{1}{2} \int \mathrm{d}^3\underline{x} \left(\sqrt{|g_{\tau_2}^{(3)}g_{\tau_2}^{\tau\tau}|} \,\varphi(\tau_2,\underline{x})(\vartheta_\tau\varphi)(\tau_2,\underline{x}) \right. \\ &\left. - \sqrt{|g_{\tau_1}^{(3)}g_{\tau_1}^{\tau\tau}|} \,\varphi(\tau_1,\underline{x})(\vartheta_\tau\varphi)(\tau_1,\underline{x}) \right), \end{split}$$

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where $g_{\tau_1}^{(3)}$ and $g_{\tau_2}^{(3)}$ denote the metric restricted to the hypersurfaces Σ_1 and Σ_2 respectively.

In terms of the boundary field configuration,

$$S_{[\tau_1,\tau_2],0}(\phi_1,\phi_2) = \frac{1}{2} \int \mathrm{d}^3\underline{x} \begin{pmatrix} \phi_1 & \phi_2 \end{pmatrix} W_{[\tau_1,\tau_2]} \begin{pmatrix} \phi_1 \\ \phi_2 \end{pmatrix},$$

where the $W_{[\tau_1,\tau_2]}$ is a 2×2 matrix with elements $W^{(i,j)}_{[\tau_1,\tau_2]}, (i,j=1,2),$ given by

$$\begin{split} & \mathcal{W}_{[\tau_{1},\tau_{2}]}^{(1,1)} = -\sqrt{|g_{\tau_{1}}^{(3)}g_{\tau_{1}}^{\tau\tau}|} \, \frac{\Delta_{1}(\tau_{1},\tau_{2})}{\Delta(\tau_{1},\tau_{2})}, \quad \mathcal{W}_{[\tau_{1},\tau_{2}]}^{(1,2)} = -\sqrt{|g_{\tau_{1}}^{(3)}g_{\tau_{1}}^{\tau\tau}|} \, \frac{\Delta_{2}(\tau_{1},\tau_{1})}{\Delta(\tau_{1},\tau_{2})}, \\ & \mathcal{W}_{[\tau_{1},\tau_{2}]}^{(2,1)} = \sqrt{|g_{\tau_{2}}^{(3)}g_{\tau_{2}}^{\tau\tau}|} \, \frac{\Delta_{1}(\tau_{2},\tau_{2})}{\Delta(\tau_{1},\tau_{2})}, \qquad \mathcal{W}_{[\tau_{1},\tau_{2}]}^{(2,2)} = \sqrt{|g_{\tau_{2}}^{(3)}g_{\tau_{2}}^{\tau\tau}|} \, \frac{\Delta_{2}(\tau_{1},\tau_{1})}{\Delta(\tau_{1},\tau_{2})}, \end{split}$$

where

$$\Delta_1(\tau_1,\tau_2) := \partial_{\tau} \Delta(\tau,\tau_2) \big|_{\tau=\tau_1} \qquad \Delta_2(\tau_1,\tau_2) := \partial_{\tau} \Delta(\tau_1,\tau) \big|_{\tau=\tau_2}.$$

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The symplectic form on the space of smooth solutions in M is

$$\omega(\varphi_1,\varphi_2) = \frac{1}{2} \int_{\Sigma} \mathrm{d}^3\underline{x} \sqrt{|g^{(3)}g^{\tau\tau}|} \left(\varphi_1 \, \vartheta_\tau \varphi_2 - \varphi_2 \, \vartheta_\tau \varphi_1\right),$$

which is independent of the choice of leaf Σ of the foliation.

Consequently the operator

$$\mathcal{W} \coloneqq \sqrt{|g_\tau^{(3)}g_\tau^{\tau\tau}|} \Delta_2(\tau,\tau) = -\sqrt{|g_\tau^{(3)}g_\tau^{\tau\tau}|} \Delta_1(\tau,\tau)$$

is independent of τ .

This implies that

$$W^{(1,2)}_{[\tau_1,\tau_2]} = W^{(2,1)}_{[\tau_1,\tau_2]}.$$

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Schrödinger-Feynman quantization

- The state space H_Σ for a hypersurface Σ is the space of functions on field configurations K_Σ on Σ.
- Inner product,

$$\langle \psi_2 | \psi_1 \rangle = \int_{\mathcal{K}_{\Sigma}} \mathcal{D} \phi \, \psi_1(\phi) \overline{\psi_2(\phi)}.$$

• Amplitude for a region $M, \psi \in \mathcal{H}_{\partial M}$,

$$\rho_{\textit{M}}(\psi) = \int_{\textit{K}_{\eth\textit{M}}} \mathcal{D}\phi\,\psi(\phi) \textit{Z}_{\textit{M}}(\phi),$$

where Z_M is the field propagator given by the Feynman path integral,

$$Z_{\boldsymbol{M}}(\boldsymbol{\varphi}) = \int_{\mathcal{K}_{\boldsymbol{M}}, \boldsymbol{\varphi}|_{\partial \boldsymbol{M}} = \boldsymbol{\varphi}} \mathcal{D}\boldsymbol{\varphi} \, e^{\mathrm{i} S_{\boldsymbol{M}}(\boldsymbol{\varphi})}, \quad \forall \boldsymbol{\varphi} \in \mathcal{K}_{\partial \boldsymbol{M}}.$$

The integral is over the space K_M of space-time field configurations ϕ in the interior of M which agree with ϕ on the boundary ∂M .

Free theory: field propagator

The field propagator in the region M results to be

$$Z_{[\tau_{1},\tau_{2}],0}(\phi_{1},\phi_{2}) = \int_{\varphi|_{\Sigma_{1,2}}=\phi_{1,2}} \mathcal{D}\varphi \ e^{iS_{[\tau_{1},\tau_{2}],0}(\varphi)}.$$

The integral is evaluated by shifting the integration variable by a classical solution matching the boundary configurations φ_1 and φ_2 at $\tau = \tau_1$, $\tau = \tau_2$ respectively,

$$Z_{[\tau_1,\tau_2],0}(\phi_1,\phi_2) = N_{[\tau_1,\tau_2],0} e^{iS_{[\tau_1,\tau_2],0}(\phi_1,\phi_2)},$$

where $S_{[\tau_1,\tau_2],0}(\phi_1,\phi_2)$ is the free action and the normalization factor is formally given by

$$N_{[\tau_1,\tau_2],0} = \int_{\varphi|_{\Sigma_{1,2}}=0} \mathcal{D}\phi \, e^{\mathrm{i}S_{[\tau_1,\tau_2],\mathbf{0}}(\phi)} = \left(\det\frac{\mathrm{i}\mathcal{W}_{[\tau_1,\tau_2]}^{(1,2)}}{2\pi}\right)^{-1/2}$$

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The field propagator satisfies the composition rule

$$Z_{[\tau_1,\tau_3],0}(\varphi_1,\varphi_3) = \int \mathcal{D}\varphi_2 \, Z_{[\tau_1,\tau_2],0}(\varphi_1,\varphi_2) Z_{[\tau_2,\tau_3],0}(\varphi_2,\varphi_3)$$

and the identity

$$\int \mathcal{D}\phi_2 \, \overline{Z_{[\tau_1,\tau_2],0}(\phi_1,\phi_2)} Z_{[\tau_1,\tau_2],0}(\tilde{\phi}_1,\phi_2) = \delta\left(\phi_1,\tilde{\phi}_1\right),$$

where φ_1 and $\tilde{\varphi}_1$ are field configurations on Σ_1 . The above identity can be interpreted in terms of the unitarity of the evolution implemented by the field propagator.

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Vacuum state

The vacuum wave functional has the form of a Gaussian,

$$\psi_{\Sigma,\mathbf{0}}(\varphi) = C_{\Sigma} \exp\left(-\frac{1}{2} \int \mathrm{d}^3 s \, \varphi(s)(A_{\Sigma} \varphi)(s)\right),$$

s being a generic coordinate system on Σ , A_{Σ} is some operator

$$A_{\Sigma} = -\mathrm{i} \sqrt{|g_{\Sigma}^3|} \, rac{\partial_n(\Upsilon(s^0))}{\Upsilon(s^0)},$$

where g_{Σ}^3 is the determinant of the 3-metric on Σ , $\partial_n = \sqrt{|g_{\Sigma}^{00}|} \partial/\partial s^0$ is the normal derivative to Σ and

$$\Upsilon(s^0) := c_a X_a(s^0) + c_b X_b(s^0),$$

where $c_{a,b}$ are complex numbers s.t. $\overline{c_a}c_b - \overline{c_b}c_a \neq 0$, and s^0 is the parameter indexing the foliation. The normalization factor is

$$C_{\Sigma} = \det\left(i\frac{\sqrt{|g_{\Sigma}^{3}g_{\Sigma}^{00}|}(\overline{c_{a}}c_{b} - \overline{c_{b}}c_{a})\Delta_{1}(s^{0}, s^{0})}{2\pi\Upsilon^{2}(s^{0})}\right)^{1/4}.$$

Coherent states are defined, in the interaction picture, in terms of a complex function ξ ,

$$\psi_{\tau,\xi}(\phi) = \mathcal{K}_{\tau,\xi} \exp\left(\int \mathrm{d}^3\underline{x}\,\frac{\xi(\underline{x})}{\Upsilon(\tau)}\,\phi(\underline{x})\right)\psi_{\tau,0}(\phi),$$

and the normalization factor results to be

$$\mathcal{K}_{\tau,\xi} = \exp\left(-\frac{1}{2}\int \mathrm{d}^3\underline{x} \, \frac{1}{\mathrm{i}\mathcal{W}_{[\tau_1,\tau_2]}^{(1,2)}(\overline{c_a}\,c_b - \overline{c_b}\,c_a)\,\Delta(\tau_1,\tau_2)} \left(\frac{\overline{\gamma(\tau)}}{\gamma(\tau)}\xi^2(\underline{x}) + \left|\xi(\underline{x})\right|^2\right)\right)$$

Free amplitude

The free amplitude for the coherent state $\psi_{\tau_1,\xi_1} \otimes \overline{\psi_{\tau_2,\xi_2}} \in \mathcal{H}_1 \otimes \mathcal{H}_2^*$ is
$$\begin{split} &\rho_{[\tau_1,\tau_2],0}(\psi_{\tau_1,\xi_1} \otimes \overline{\psi_{\tau_2,\xi_2}}) \\ &= \int \mathcal{D}\phi_1 \, \mathcal{D}\phi_2 \, \overline{\psi_{\tau_2,\xi_2}(\phi_2)} \, \psi_{\tau_1,\xi_1}(\phi_1) \, Z_{[\tau_1,\tau_2],0}(\phi_1,\phi_2), \\ &= \exp\left(-\frac{1}{2} \int d^3x \, \frac{1}{iW_{[\tau_1,\tau_2]}^{(1,2)} \left(\overline{c_a} \, c_b - \overline{c_b} \, c_a\right) \, \Delta(\tau_1,\tau_2)} \left(|\xi_1(\underline{x})|^2 + |\xi_2(\underline{x})|^2 - 2\overline{\xi_2(\underline{x})} \, \xi_1(\underline{x})\right)\right). \end{split}$$

This amplitude does not depend on τ_1 and τ_2 , since the product $W^{(1,2)}_{[\tau_1,\tau_2]}\Delta(\tau_1,\tau_2)$ does not depend on τ_1 and τ_2 . This was expected since we are considering the free evolution of states in the interaction picture.

Interacting theory

Consider the interaction of the scalar field with a real source field $\boldsymbol{\mu}$ described by the action

$$S_{M,\mu}(\phi) = S_{M,0}(\phi) + \int_M \mathrm{d}^4 x \sqrt{-g(x)} \,\mu(x) \,\phi(x).$$

We assume that the field μ is confined in the interior of the region M, i.e. $\mu(x) = 0$ for $x \in \partial M$ and $x \notin M$. In the region $M = [\tau_1, \tau_2] \times \mathbb{R}^3$ the action reads

$$\mathcal{S}_{[\tau_1,\tau_2],\mu}(\phi_1,\phi_2) = \mathcal{S}_{[\tau_1,\tau_2],0}(\phi_1,\phi_2) + \int \mathrm{d}^3\underline{x} \big(\mu_1(\underline{x}) \ \phi_1(\underline{x}) + \mu_2(\underline{x}) \ \phi_2(\underline{x})\big),$$

where

$$\begin{split} \mu_1(\underline{x}) &:= \int_{\tau_1}^{\tau_2} \mathrm{d}\tau \, \sqrt{-g(\tau,\underline{x})} \, \frac{\Delta(\tau,\tau_2)}{\Delta(\tau_1,\tau_2)} \, \mu(\tau,\underline{x}), \\ \mu_2(\underline{x}) &:= \int_{\tau_1}^{\tau_2} \mathrm{d}\tau \, \sqrt{-g(\tau,\underline{x})} \, \frac{\Delta(\tau_1,\tau)}{\Delta(\tau_1,\tau_2)} \, \mu(\tau,\underline{x}). \end{split}$$

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Field propagator

The field propagator can be expressed in terms of the free one as

$$\begin{split} Z_{[\tau_1,\tau_2],\mu}(\phi) &= Z_{[\tau_1,\tau_2],0}(\phi_1,\phi_2) \\ &\times \frac{\mathcal{N}_{[\tau_1,\tau_2],\mu}}{\mathcal{N}_{[\tau_1,\tau_2],0}} \, \exp\left(\mathrm{i} \int \mathrm{d}^3\underline{x} \left(\mu_1(\underline{x}) \, \phi_1(\underline{x}) + \mu_2(\underline{x}) \, \phi_2(\underline{x})\right)\right), \end{split}$$

The normalization factor $N_{[\tau_1,\tau_2],\mu}$ is formally equal to

$$N_{[\tau_1,\tau_2],\mu} = \int_{\varphi|_{\tau_{1,2}}=0} \mathcal{D}\varphi \, e^{\mathrm{i}S_{[\tau_1,\tau_2],\mu}(\varphi)} = N_{[\tau_1,\tau_2],0} \exp\left(\frac{\mathrm{i}}{2} \int \mathrm{d}^4 x \, \sqrt{-g(x)} \, \alpha(x) \, \mu(x)\right).$$

The integral has been evaluated by shifting the integration variable by a solution α of the inhomogeneous Klein-Gordon equation, $(\Box + m^2) \alpha(x) = \mu(x)$, with vanishing boundary conditions, $\alpha(\tau_1, \underline{x}) = \alpha(\tau_2, \underline{x}) = 0$. α can be written as

$$\begin{split} \alpha(\mathbf{x}) = \int_{\tau_1}^{\tau_2} \mathrm{d}\tau' \sqrt{-g(\tau',\underline{\mathbf{x}})} \left(\theta(\tau'-\tau) \frac{\Delta(\tau_1,\tau)\Delta(\tau',\tau_2)}{W_{[\tau_1,\tau_2]}^{(1,2)}\Delta^2(\tau_1,\tau_2)} \right. \\ \left. + \theta(\tau-\tau') \frac{\Delta(\tau_1,\tau')\Delta(\tau,\tau_2)}{W_{[\tau_1,\tau_2]}^{(1,2)}\Delta^2(\tau_1,\tau_2)} \right) \mu(\tau',\underline{\mathbf{x}}). \end{split}$$

It can be shown that the field propagator satisfies the identity

$$\int \mathcal{D}\phi_2 \, \overline{Z_{[\tau_1,\tau_2],\mu}(\phi_1,\phi_2)} Z_{[\tau_1,\tau_2],\mu}(\tilde{\phi}_1,\phi_2) = \delta\left(\phi_1,\tilde{\phi}_1\right),$$

representing the unitarity of the quantum evolution implemented by the field propagator $Z_{[\tau_1,\tau_2],\mu}.$

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Amplitude

The amplitude for the boundary state $\psi_{\tau_1,\xi_1}\otimes\overline{\psi_{\tau_2,\xi_2}}$ in the presence of the source field μ is

$$\rho_{[\tau_1,\tau_2],\mu}(\psi_{\tau_1,\xi_1}\otimes\overline{\psi_{\tau_2,\xi_2}}) = \int \mathcal{D}\phi_1 \mathcal{D}\phi_2 \overline{\psi_{\tau_2,\xi_2}(\phi_2)} \psi_{\tau_1,\xi_1}(\phi_1) Z_{[\tau_1,\tau_2],\mu}(\phi_1,\phi_2),$$

and can be expressed in terms of the free amplitude

where the function $\hat{\boldsymbol{\xi}}$ is

$$\hat{\xi}(x) = \frac{1}{W_{[\tau_{1},\tau_{2}]}^{(1,2)}\left(\overline{c_{a}} c_{b} - \overline{c_{b}} c_{a}\right) \Delta(\tau_{1},\tau_{2})} \left(\left(\overline{\Upsilon(\tau)}\xi_{1}\right)(\underline{x}) + \left(\Upsilon(\tau)\overline{\xi_{2}}\right)(\underline{x})\right),$$

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and G_F is given by

$$\begin{split} \mathcal{G}_{\mathcal{F}}(\mathbf{x},\mathbf{x}') = & \int \mathrm{d}^{3}\underline{k} \, \frac{1}{W_{[\tau_{1},\tau_{2}]}^{(1,2)} \, \Delta(\tau_{1},\tau_{2})(\overline{c_{a}} \, c_{b} - \overline{c_{b}} \, c_{a})} \\ & \left(\theta(\tau'-\tau) \left(\Upsilon(\tau)\varphi_{\underline{k}}\right)(\underline{\mathbf{x}}) \, \overline{\left(\Upsilon(\tau')\varphi_{\underline{k}}\right)}(\underline{\mathbf{x}}') + \theta(\tau-\tau') \left(\Upsilon(\tau')\varphi_{\underline{k}}\right)(\underline{\mathbf{x}}') \, \overline{\left(\Upsilon(\tau)\varphi_{\underline{k}}\right)}(\underline{\mathbf{x}}) \right) \end{split}$$

In Minkowski spacetime, in a slice region where τ represents the Minkowski time, G_F coincides with the standard Feynman propagator. The same happens in de Sitter and Rindler spaces with τ equal to de Sitter conformal time and Rindler time respectively. G_F plays the rôle of the **Feynman propagator** for the scalar field theory

 G_F plays the role of the **reginman propagator** for the scalar field theory defined in the slice region and satisfies the inhomogeneous Klein-Gordon equation in both variables x and x'

$$(\Box + m^2) G_F(x, x') = (-g(x))^{-1/2} \delta^4(x - x').$$

This expression is independent of τ_1 and τ_2 . The limit of the amplitude for asymptotic values of τ_1 and τ_2 is trivial, and we can interpret it as the S-matrix for the scalar theory in the presence of a source field.

General interaction

Consider the general interacting theory

$$S_{M,V}(\phi) = S_{M,0}(\phi) + \int_M \mathrm{d}^4 x \, V(x,\phi(x)).$$

where V is an arbitrary potential. The exponential of i times this action can be written as

$$e^{\mathrm{i}S_{\boldsymbol{M},\boldsymbol{\nu}}(\boldsymbol{\Phi})} = \exp\left(\mathrm{i}\int_{\boldsymbol{M}} \mathrm{d}^{4}x\,\sqrt{-g(x)}\,V\left(x,-\mathrm{i}\frac{\delta}{\delta\mu(x)}\right)\right)e^{\mathrm{i}S_{\boldsymbol{M},\mu}(\boldsymbol{\Phi})}\bigg|_{\mu=0}\,,$$

where $S_{M,\mu}$ is the action in the presence of a source interaction. We assume that the potential V vanishes outside the region M. The corresponding field propagator is

$$Z_{M,V}(\varphi) = \exp\left(i\int_{M} d^{4}x \sqrt{-g(x)} V\left(x, -i\frac{\delta}{\delta\mu(x)}\right)\right) Z_{M,\mu}(\varphi)\Big|_{\mu=0},$$

and the amplitude for the general interacting theory is

$$\rho_{M,V}(\psi) = \exp\left(i\int_{M} \mathrm{d}^{4}x \sqrt{-g(x)} V\left(x, -i\frac{\delta}{\delta\mu(x)}\right)\right) \rho_{M,\mu}(\psi) \bigg|_{\mu=0}.$$