

Exactly Solvable LQC: New Insights on Some Old Questions

(Work with Abhay Ashtekar and Alejandro Corichi (*To appear*))

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Motivation

- Extensive analytical and numerical methods in LQC: valuable insights on singularity resolution in symmetry reduced models. **Big Bang** replaced by **Big Bounce**. Non-local character of curvature leads to repulsive QG effects at Planck scale \Rightarrow significant departures from classical dynamics. At low curvatures: LQC \rightarrow GR

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 - What happens to the singularity resolution when "area gap goes to zero"? Does LQC approach WDW in this limit?
 - Does a continuum limit of LQC exist when area gap $\rightarrow 0$? Or is LQC a fundamentally discrete theory?

Outline

- A very short introduction of LQC
- Simplified LQC (SLQC) (contrast to Bojowald's model) and WDW in b representation
- Volume observable and its fluctuations (for arbitrary states)
- Comparison of SLQC and WDW
- Issues of limit

LQC: A Very Brief Introduction

- Gravitational + Matter Constraint (Massless Scalar)

$$-\gamma^{-2} \int_{\mathcal{V}} d^3x \varepsilon_{ijk} \frac{E^{ai} E^{bj}}{\sqrt{|\det E|}} F_{ab}^k + 8\pi G \frac{P_\phi^2}{2\sqrt{|\det E|}} = 0$$

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- Quantum constraint: $(\sin(\lambda b) \hat{A} \sin(\lambda b) + KB(\nu) \hat{P}_\phi^2) \psi(\nu, \phi) = 0$

$$b := c/|p|^{1/2}, \quad \{b, \nu\} = 2, \quad \nu = V/(2\pi\gamma\ell_{\text{P}}^2), \quad \lambda^2 = \text{Area Gap.}$$

$$\hat{b}\psi(\nu) = -2i\frac{\partial}{\partial\nu}\psi(\nu), \quad \exp(i\lambda b/2)\psi(\nu) = \psi(\nu + \lambda)$$

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- For $\nu > 1$, $A(\nu) = 1$ and $B(\nu) \rightarrow 1/|\nu|$. States corresponding to a large classical universe at late times do not see $\nu \leq 1$.

Simplified LQC

- Quantum Constraint in b representation:

$$\Theta(\mathbf{b})\chi(\mathbf{b}, \phi) = -12\pi G \frac{\sin(\lambda\mathbf{b})}{\lambda} \frac{\partial}{\partial\mathbf{b}} \frac{\sin(\lambda\mathbf{b})}{\lambda} \frac{\partial}{\partial\mathbf{b}} \chi(\mathbf{b}, \phi) = -\partial_\phi^2 \chi(\mathbf{b}, \phi)$$

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- Introduce $x := (12\pi G)^{-1/2} \log(\tan(\lambda\mathbf{b}/2)) \Rightarrow \partial_\phi^2 \chi = \partial_x^2 \chi$

General solution:

$$\chi = \chi_+(\phi + x) + \chi_-(\phi - x) := \chi_+(x_+) + \chi_-(x_-)$$

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- Physical states anti-symmetric in \mathbf{b} : $\chi(\mathbf{b}, \phi) = -\chi(\mathbf{b} - \pi/2, \phi)$. Imposes relation between χ_+ and χ_-

Wheeler-DeWitt Theory

- Quantum constraint in b representation:

$$\Theta(b)\chi(b, \phi) = -12\pi G b \frac{\partial}{\partial b} b \frac{\partial}{\partial b} \chi(b, \phi) = -\partial_\phi^2 \chi(b, \phi)$$

- As in SLQC, we have an internal clock, Physical inner product and Dirac Observables.

- Introduce

$$y := (12\pi G)^{-1/2} \log (b/2b_o)$$

\Rightarrow

$$\partial_\phi^2 \chi(\phi, y) = \partial_y^2 \chi(\phi, y)$$

- General solution:

$$\chi = \chi_+(\phi + y) + \chi_-(\phi - y) := \chi_+(y_+) + \chi_-(y_-)$$

- Unlike SLQC, χ_+ (expanding) and χ_- (contracting) are disjoint.

Volume observable in WDW

$$\begin{aligned}(\chi, \hat{V} |_{\phi} \chi)_{\text{phy}} &= 2\pi\gamma\ell_{\text{P}}^2 (\hat{\nu}\chi, \hat{\nu}\chi)_{\text{kin}} \\ &= \frac{16\gamma\ell_{\text{P}}^2}{\sqrt{12\pi G} b_o^2} \int_{-\infty}^{\infty} dy_+ \left| \frac{d\chi_+}{dy_+} \right|^2 e^{\sqrt{12\pi G}(\phi - y_+)} \\ &= V_o e^{\sqrt{12\pi G}\phi} .\end{aligned}$$

- As $\phi \rightarrow -\infty$, $\langle \hat{V} |_{\phi} \rangle \rightarrow 0$. The backward evolution leads to the big bang singularity.
- Fluctuations:

$$(\chi, \hat{V}^2 |_{\phi} \chi)_{\text{phy}} = W_0 e^{2\sqrt{12\pi G}\phi}$$

$$((\Delta V |_{\phi_0}) / \langle \hat{V} |_{\phi_0} \rangle)^2 = (W_0 / V_0)^2 - 1 .$$

Remains constant with evolution.

Volume observable in SLQC

$$\begin{aligned}(\chi, \hat{V} |_{\phi} \chi)_{\text{phy}} &= \frac{8\gamma \ell_{\text{P}}^2 \lambda^2}{\sqrt{12\pi G}} \left[\int_{-\infty}^{\infty} dx_+ \left| \frac{d\chi_+}{dx_+} \right|^2 \cosh(\sqrt{12\pi G}(x_+ - \phi)) \right. \\ &\quad \left. + \int_{-\infty}^{\infty} dx_- \left| \frac{d\chi_-}{dx_-} \right|^2 \cosh(\sqrt{12\pi G}(-x_- + \phi)) \right] \\ &= I_1 e^{-\sqrt{12\pi G}\phi} + I_2 e^{\sqrt{12\pi G}\phi}\end{aligned}$$

● There exists a minimum value of $\langle V |_{(\phi=\phi_B)} \rangle$ which occurs at

$$\phi_B = (2\sqrt{12\pi G})^{-1} \log(I_1/I_2)$$

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- $\langle V |_{\phi} \rangle$ is symmetric around the bounce point.

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- Relative dispersion bounded in time evolution. A single condition on the infinite dimensional space of initial data: $J_1 = \frac{I_1^2}{I_2^2} J_2$ implies symmetric fluctuations around bounce point and equality of asymptotic values.

Comparison between WDW and SLQC

- For a fixed value of λ select $\Psi_0(\mathbf{b})$. Initially, for $\phi = 0$

$$\langle \hat{V} \rangle_\lambda(\phi = 0) = \langle \hat{V} \rangle_{\text{WDW}}(\phi = 0) =: V_0$$

Difference between expectation values:

$$\langle \hat{V} \rangle_{\text{WDW}}(\phi) - \langle \hat{V} \rangle_\lambda(\phi) = (V_0 - I_2) e^{\sqrt{12\pi G} \phi} - I_1 e^{-\sqrt{12\pi G} \phi}$$

Relative difference: bounded in future evolution

$$|\langle \hat{V} \rangle_{\text{WDW}}(\phi) - \langle \hat{V} \rangle_\lambda(\phi)| / \langle \hat{V} \rangle_{\text{WDW}}(\phi) \leq \delta := I_1 / V_0 \text{ (very small)}$$

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- For a given ϕ_T and $\epsilon > 0$, $\exists \lambda_{(\epsilon, T)} > 0$ such that,

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- For any $N > 0$ (arbitrarily large) $\exists \phi$ such that

$$|\langle \hat{V} \rangle_{\text{WDW}}(\phi) - \langle \hat{V} \rangle_\lambda(\phi)| > N$$

Fundamental discreteness of SLQC

- Start with an arbitrary λ_o , refine the area gap ($\lambda_o \rightarrow \lambda$).
For $\lambda < \lambda_o$, $\chi_i \in \mathcal{H}_{\lambda_o}$ under embedding $\chi_i \in \mathcal{H}_{\lambda}$.
Under renormalization $\chi^\lambda := \sqrt{\lambda_o/\lambda} \chi^{\lambda_o}$, $|\chi^\lambda|^2 = |\chi^{\lambda_o}|^2$.

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- Refinement in $\lambda \Rightarrow x_\lambda \neq x_{\lambda_o} \Rightarrow I_{1,2}(\lambda) \neq I_{1,2}(\lambda_o)$. $I_2(\lambda)$ is a monotonic decreasing function. As $\lambda \rightarrow 0$, $I_2(\lambda)$ grows.

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- Consequence: In the backward evolution of an expanding branch

$$\langle \hat{V} \rangle_{\lambda_o} - \langle \hat{V} \rangle_\lambda = (I_2(\lambda_o) - I_2(\lambda)) e^{-\sqrt{12\pi G} \phi}$$

which diverges as $\phi \rightarrow -\infty$.

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$$\phi_B = (2\sqrt{12\pi G})^{-1} \log(I_1/I_2) \longrightarrow -\infty \quad \text{as} \quad \lambda \rightarrow 0.$$

- Uniform limit does not exist. Contrast with results on Harmonic Oscillator (Corichi, Vukasinac, Zapata (07)).

Summary

- Without assuming semi-classicality at any level we presented analysis of arbitrary states in SLQC and WDW. Full analytical control to make predictions from the quantum theory.
- Bounce not restricted to states which are semi-classical at late times. There is a pre-big bang branch for a dense subspace of \mathcal{H}_{phy} . Only one condition on the infinite dimensional space of initial data leads to symmetric fluctuations across the bounce point.
- SLQC and WDW approach GR at low curvatures. At large curvatures they depart significantly.
- In the backward evolution of the expanding branch for any given fixed time interval, SLQC and WDW agree to arbitrary accuracy by a choice of λ . However, for any given choice of λ , they diverge if one waits long enough.
- There is no limiting theory of SLQC when $\lambda \rightarrow 0$. Two different λ SLQC's depart in a similar way as they do from WDW. SLQC is a fundamentally discrete theory.