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Black hole radiation spectrum

Loop Quantum Gravity

in



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Jacobo Díaz-Polo

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Here: Some different new results.

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BLACK HOLES IN LQG

[A.Ashtekar, J.Baez, A.Corichi, K.Krasnov]

Inner boundary introduced in the spacetime manifold before quantization procedure (*effective treatment*)

Isolated horizon boundary conditions

Hilbert space splitting: $H_v \otimes H_s$

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Hilbert space splitting: $H_v \otimes H_s$

Surface states: U(1) Chern-Simons theory over a punctured sphere

Spin network pierces the horizon at punctures

Punctures carry some quantum labels satisfying

$$\sum_{i} m_{i} = 0$$
 (projection constraint)

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Count all possible combinations of labels

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Result:
$$S = \frac{1}{4}A - \frac{1}{2}\ln A$$
 up to γ .

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COMPUTATIONAL COUNTING

A.Corichi, E.Fernandez-Borja, J.D. [Class. Quantum Grav. 24, 243 (2007)]

In order to make an exact counting we use an explicit enumeration algorithm

Computer generates all possible combinations of labels and counts only those satisfying the required conditions

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ABCK framework formulated for large area limit

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But linear dependence and logarithmic correction are obtained \Rightarrow some confidence

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$$\Delta A = \chi \gamma \ell_P^2$$

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Where this behaviour comes from?

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Structure of the horizon area eigenstate degeneracy when introducing the "projection constraint"

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Structure with "bands" or peaks of degeneracy

Degeneracy at the peaks several orders of magnitude higher than the rest of the spectrum

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- Implications in spectroscopy

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Structure with "bands" or peaks of degeneracy

Degeneracy at the peaks several orders of magnitude higher than the rest of the spectrum

With this: - Explanation for the staircase

- Implications in spectroscopy

Consider all possible transitions regardless of the concrete quantum configuration (unknown dynamics)



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E.Fernandez-Borja, J.D. [arXiv: 0706.1979 [gr-qc]]

First approximation: consider only states at the peaks

Equidistant area spectrum with a gap $\Delta A = \chi \gamma \ell_P^2$

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Analogous to Bekenstein-Mukhanov picture

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- Statistically preferred frequencies
- Equidistant "brighter" lines sticking out from a continuous background spectrum
- The bands are not narrow \Rightarrow Broadening of the lines

CONCLUSIONS

We make an spectroscopical analysis at the kinematical level (we don't know dynamics)

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Observable qualitative effects in the spectrum

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Considering all possible transitions, IH introduce a quantum gravity imprint in the black hole radiation spectrum

Observable qualitative effects in the spectrum

 $ω_0$ proportional to $γ \Rightarrow$ hypothetical observational determination (GRB's ???)

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THE END